GW151226: FACTSHEET

BACKGROUND IMAGES: TIME-FREQUENCY TRACE (TOP) AND SIGNAL-TO-NOISE RATIO TIME-SERIES (BOTTOM) IN THE TWO LIGO DETECTORS; EXAMPLE WAVEFORM (MIDDLE)

observed by	LIGO L1, H1	duration from 35 Hz	~1 s
source type	black hole (BH) binary	# cycles from 35 Hz	~55
date	26 Dec 2015	signal arrival time	arrived in H1 1 ms after
time	03:38:53 UTC	delay	L1
distance	250 to 620 Mpc	peak GW strain	~ 3.4 x 10 ⁻²²
redshift	0.05 to 0.13	peak displacement of interferometers arms	~ ±0.7 am
signal-to-noise ratio	13		_ 500 500
false alarm prob.	~ 1 in 10 million	frequency/wavelength at peak GW strain	420 Hz, 710 km
Source Mass	ses M⊙	peak speed of BHs	~ 0.6 c
total mass	20 to 28	\ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \	2 to 4 x 10 ⁵⁶ erg s ⁻¹
primary BH	11 to 23	peak GW luminosity	
secondary BH	5 to 10	radiated GW energy	0.8-1.1 M⊙
	(template) 9 to 27 0.48	_{0.4} remnant ringdown fro	eq. ~ 750 Hz
mass ratio	> 0.28	remnant damping tin	ne 0.00 ~ 1.30ms
spin of one of the	> 0.2	remnant size, area	60 km, 3.5 x 10 ⁴ km ²
black holes remnant BH spin	0.7 to 0.8	online trigger latency	~ 67 s
resolved to	~850 sq. deg.	# offline analysis pipeli	nes 2

Parameter ranges correspond to 90% credible bounds. **Acronyms**: L1/H1=LIGO Livingston/Hanford; Mpc=mega parsec=3.2 million lightyear, am=attometer=10⁻¹⁸ m, M⊙=1 solar mass=2 x 10³⁰ kg

